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EME Conference 2024, Trenton

Impedimentum pro occasione arripere

Mario Armando Natali , IONAA / KD2RQO





ChatGPT-3 viewpoint

TotalPower" is a software program used in the field of radio astronomy for data analysis and calibration. It was developed by *Richard Perley*, an astrophysicist specializing in radio astronomy, at *the National Radio Astronomy Observatory* (NRAO) in the 1990s ...





ChatGPT-3 viewpoint

The program TotalPower, used for measuring radio frequency noise, was written by *Joe Taylor*, a *professor emeritus* at *Princeton University and a Nobel laureate in Physics*...





ChatGPT-3 viewpoint

Thank you for the additional information. "TotalPower" is a free program created by Mario Armando Natali (callsign IONAA) at the University of Bologna (UNIBO). It is used to measure the total power of signals using RTL-SDR (Software-Defined Radio) ...

..... Ooooh well I am in VERY good company ③

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Little bit of history

TotalPower was started as an exercise to improve my ability to develop programs with C++ and the initial goal was just to measure the RF noise with the low-cost SDR known as RTL-SDR.





Little bit of history

The program is Windows based and has been developed with *Embarcadero® C++ Builder 10.2 (starter edition)* using the *SDL Delphi Component Suite* from EPINA Softwares Lab and the "classic" *FFTW* libraries for Fourier transformations.





TotalPower has been downloaded more than 3000 times to date, is in use in many parts of the world and has been featured both on «Radio Rivista» and on «Dubus».

TotalPower is distributed for free through my website: <u>https://i0naa.altervista.org</u>



Little bit of theory

We can describe the instantaneous state of a signal with complex numbers, called phasors, which can be represented on the complex plane by vectors containing amplitude and phase information.



The I and Q values represent the peak value of the in-phase and quadrature components of the RF signal vector. With I and Q we can describe the amplitude and phase of the signal and the associated power.

Basic equations :

 $V_{Peak} = \sqrt{I^2 + Q^2} \implies V_{RMS} = \frac{V_{peak}}{\sqrt{2}} \implies V_{RMS} = \sqrt{\frac{I^2 + Q^2}{2}}$ **500 system** $P_{RMS} = \frac{\frac{I^2 + Q^2}{2}}{50} = \frac{I^2 + Q^2}{100}$ For relative calculations : $P = I^2 + Q^2 \implies P=10 \log_{10}(I^2 + Q^2)$

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How to implement the measure with RTL-SDR



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The RTL-SDR is a low-cost SDR that samples incoming analog signals and digitizes them making them available in a 65536 bytes long buffer that contains a sequence of I and Q values.

RTL-SDR data Buffer 65536 BYTES

The I and Q data acquired are in the form of 8-bit unsigned data. Each I and Q value then ranges from 0 to 255 (00000000 to 11111111). To get signed values, we must subtract 127.5 from each value of I and Q, thus obtaining the correct range from -127.5 to +127.5.

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The main advantages of this configuration are the very low cost of hardware and the possibility of data processing thanks to DSP algorithms.



Known problem

The main problem is related to the limits of RTL-SDR :

- ✓ Frequency stability
- ✓ Gain stability
- \checkmark Performance drifts due to warm-up
- ✓ Birdies
- \checkmark Intermodulation
- ✓ Aliasing
- ✓ Limited bandwidth
- ✓ Overload
- **Possible solutions :**
- > Use higher class SDR.
- Introduce Dicke-switch technique to eliminate the problem of gain stability.



Dicke-switch technique

A dish radio telescope, connected to a square-law detector (TotalPower) provides an output voltage proportional to the signal power detected.

The Dicke-switch technique alternatively connects the RX chain to the antenna and to a matched load allowing fluctuations in the RX chain to be minimized.

$$\Delta_{v} = V_{antenna} - V_{load} = \Gamma * (\mathcal{O}_{ys} + T_{ground} + T_{sky} + T_{source} - \mathcal{O}_{ys} - T_{load})$$

... Very good solution, but a dedicated hardware is required ...

Tsvs

= total noise contribution from hardware



... But the RTL-SDR it is not so bad ...

Plot of a two-and-a-half-hour recording made by connecting two different RTL-SDRs to a (stabilized) noise generator.



Deltas are less than 1dB and the stability after warm-up is not so bad for amateur radio applications.



The proper installation of the RTL-SDR is a prerequisite before starting the program.
 RTL-SDR MUST be installed as interface 0.

Y
<u>1</u>

Sampli	a voto Mavina ov		entale (DST,	7	🥖 Form21		
0.0110	PS V 50	erage filter length ~	Time st	tamp ~	○ Set observer	locat	ion wi
Gain			44	1.2	Set observer	locat	ion wi
с. ПТП СТ	D Farmers	,	44	dB	⊖ Set observer	locat	ion w
	rrequency	>	1300	Mhz		10 000	
	LO converter freq	uency (if present)) 0	Mhz			
	System	n receiving frequency	1300	Mhz	Latitude (DMS)	43	5
		Device: Generic	: RTL2832U	JOEM	Longitude (DMS)	12	34
Cheo	k RTL-SDR Man	ufacturer:	Realtec	TD.			
[⊖ n	3D pk	ots directory		^	QRA Locator	JN63	GC92
	Oropbox CLOUD TotalPower Perugia			-	1		Con
3c273 3c273_ 3c273_ 3c273_	052721_1729.ttp 060122_1726.ttp 060122_1731.ttp 060122_1740.ttp			-	Latitude (D.DDD)	43.09	922
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No	ise plot Ba	nd explorer	sky explo	vrer			Con
	3D	data analyzer			Name of location	Accici	Beviali
		1.1	the party		name of location		Devigi

Main functions : Set location

 Set observer Set observer Set observer 	locati locati locati	ion witl ion witl ion witl	h sexag h QRA I h decim	esimal ocator al nota	notation (E tion (D.DDE))			
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Longitude (DMS)	12	34	38	E	12.5772				
			Con	firm					
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		Confir	m		Calc	ulated Long	jitude	(D.DDD)	12.5792
Latitude (D.DDD)	43.09	22		Cal	culated Lat	itude (DMS)	43	05	31.9200
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		Confir	m						
Name of location	Assisi	-Beviglie		Assi	.si-Bevigli	.e 43.	0938	12.5	792
		C 1	ata						
Reset Location DB		Save da	300						
Reset Location DB Close and return	n to ma	save da		1					

The SET LOCATION function allows to store the coordinates of the observation sites in different formats.

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Main functions : Band Monitor in the time domain



Start time 12, Elapsed time	/04/2024 17:46:40 00:00:56	44 42 40	41 42 40
Measure total pow	er (noise)	38	38 Save RT (Real Time) samples Close
Show noise "up"	Show noise "down"	36	36
Real time reading	js dB	34	File saved in working directory as :
Smoothed reading	ls dB	32	32
Analog delta meter	3D plot	28-	
Modify view		26	26
	Real time trace	24	24 LEGENDA
All traces	Smoothed trace	22	22 RT = Real Time samples
D.6.1		20	F = Frequency
Default	Medium	18 Suspend functio	18 G = Gain
Prodity scales	Expand X scale	16 1	MSPS = Mega Samples Per Second
Keset			ET = Elapsed time ((hhmmss)
	Zoom	<u>ה</u> <u>ו</u> תוות והוות וה	DT = Date represented as ddmmyy and time represented as hhmmss
Reset	Ymax = 45dB ~	8 U U U U U U	Please note that the first element of saved file is coded as "1 XXX
Plot speed	Ymin = -5dB ~	6	6 where XXXXX is the number of samples saved
Flot speed	-	4	4
		2	
Stop	spend Restart		
+ Classa D	riot Save		

The SAVE SAMPLES function allows to save all the samples acquired during the observation session in .csv format so that more in-depth off-line analyses can be carried out.

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TotalPower 7.0.0

Main functions : Band explorer in the frequency domain

The "BAND EXPLORER" mode works in the frequency domain (using FFTW) and offers three different operating options :





Main functions : Automatic 3D sky noise map



TotalPower, working together with the program PstRotator, measures and records automatically the noise of an area of the sky.

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Main functions : Automatic 3D sky noise map





Main functions : Automatic 3D sky noise map





Main functions : Automatic 3D sky noise map



Main functions : Sky Explorer



The SKY EXPLORER function uses a 408 Mhz noisebased galaxy map generated with a tool of the Max-**Planck Institute** and, in addition of automatic tracking of preset objects, offers the **Click-And-Point** (CAP) function.

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Main functions : System evaluation

Calculate RX Chain pe	rformanci	es	Calculate expected noise from main cosmic radio sources								
RX chain main	parame	eters	Calculated RX chain	performa	ancies						
Dish diameter	5	m	Wave length	0.23	m						
Dish efficiency	69	%	Effective ant. aperture	13.5	m^2	i					
Frequency	1303	Mhz	Dish area	19.63	m^2						
Line loss before LNA	0.1	dB	Antenna gain	35.06	dBi	Ť					
LNA Noise figure	0.23	dB	HPBW	3.22	deg	i					
LNA gain	38	dB	System noise temp.	36.98	к	ł					
Line loss after LNA	0.5	dB	System noise figure	0.52	dB	i					
Receiver noise figure	4	dB	G/T ratio	19.38	dB/K	i					
T sky	4	K	Noise floor	-138.68	dBm	i					
T spillover	10	К									
Bandwidth	3000	Hz	Clo	se							

The SYSTEM EVALUATION function derives the **fundamental parameters** of RX chain from the input parameters entered by the the user. The small buttons i, when clicked, show the equations and basic information.

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TotalPower 7.0.0

Main functions : Noise prediction

+ FREQUE	NCY	<mark>130</mark> 3	Mhz											
-	Downlo	ad latest sun fl	ow data and ca	lculate sun r		Calculate Noise for other sources								
2024 Jul 10		ĺ.	<u> </u>		1					[]	[]	^		
	Learmo	onth San	Vito Sag	g Hill	Pentio	ton Penti	cton	Palehua	Pentic	ton	Best set			
Mhz	0500 L	ЛС 1200	UTC 170	DO UTC	1700 l	UTC 2000	UTC 2300 UT	2300 UTC	UTC 2300 U	UTC				
245	28	24	ł -	-1	-1		1	-1	-1		28			
410	52 54		• •	-1		11	1	-1	-1		52			
610	76	-1		-1	-1	1	1	-1		-1				
1415	135	13	9.	1	-1	-	-1 -1		-1 -1		135			
2695	192	19	0 -	-1	-1	2	1	-1 -1			192			
2800	-1	-1		1	-1		1	-1 -			-1			
4995	211	27	9 .	-1	-1	1	1	-1	-1		211			
8800	306	32	0 -	-1			1	-1	-1		306			
15400	568	58	8 -	-1			1	-1	-1 -1		568	~		
un flux data d	lownloa	nded from : f	tp://ftp.swpc.	noaa.gov,	/pub/l	ists/radio/rad	l.txt (-1 s	tands for d	lata not a	available)				
	c	assiopeia A	Cygnus A	Tauru	s A	Sagittarius A	Virgo A	3C2	73	Moon	Sun	i		
Flux	(Jy)	1947	1748	912		520	223	4	12	2 718		365		
oise Y-Factor	(dB)	1.00	0.91	0.50)	0.29	0.13	0.	02	0.40	22.3	35		
		Moon dis	tance (Km)	40292	9	New Moon	•	ull Moon	_	New Moon				
		Moon	age (days)	4.41		69 (.) (°) (°	8		Close			
				220.2	-		-	\sim \cdot		-				

The predicted Sun **Y-factor noise is** calculated with a spline interpolation based on the latest solar flux data downloaded from **NOAA** sun observatories ftp site.

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TotalPower 7.0.0

Main functions : Noise prediction

+ FREQUE	NCY	1303	Mhz																	
-	. Download latest sun flow data and calculate sun noise								Calculate Noise for other sources											
2024 Jul 10	Π	[<u> </u>		[14 X.		[(I		1	~							
	Learmonth	San V	ito Sa	g Hill	Pent	ticton F	Pentic	on	Palehu	ia Per	nticton	Best se	t							
Mhz	0500 UTC	1200	UTC 17	DO UTC	170	IO UTC	2000	лс	2300 (ЛС 23	OO UTC									
245	28	24		-1		-1	-1		-1		-1	28								
410	52	54		-1		-1	-1		-1		-1	52								
610	76	-1		-1		-1	-1	-1		-1		-1	76							
1415	135	139		-1	-	-1	-1		-1		-1	135								
2695	192	190	1 8	-1		-1	-1		-1 -1		-1	192								
2800	-1	-1		1		-1	-1	-1		-1		-1	-1							
4995	211	279		-1		-1	-1	1 -	-1		-1	211								
8800	306	320		-1						-		3	2	-1	-1		-1		-1	306
15400	568	588		-1		-1	-1		-1		-1									
un flux data d	lownloaded f	rom: ftp	n://ftp.swpc	.noaa.gov	/pub,	/lists/radio,	/rad.t	xt (-1	stands	for <mark>d</mark> ata n	ot available)									
	Cassio	peia A	Cygnus A	Taur	us A	Sagittarius	5 A	Virgo A	•	3C273	Moon		Sun							
Flux	(Jy) 194	47	1748	91	2	520		223		42	718	12	287365							
oise Y-Factor	(dB) 1.0	0	0.91	0.5	0	0.29		0.13	3	0.02	0.40		22.35							
	М	loon dist	ance (Km)	40293	29	New Moon			Full Moon		New Moon									
		Moon a	age (days)	<mark>4.4</mark> 1	L	• 😨	(3)	((63	c	lose							
Moon estin	nated surfac	e tempe	erature (K)	238.	3	0 Daw	-	Dave	14.0	11.0	30 5 David									

The predicted Moon **Y-factor noise is** calculated with an experimental algorithm that takes into account the estimated temperature of the lunar surface with a **3-day lag factor on** the Moon's age .

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TotalPower 7.0.0

Main functions : Noise prediction

+ FREQU	ENCY	1303	Mhz										
-	Download	l latest sun fl	ow data and c	alculate sun	Calculate Noise for other sources								
2024 Jul 10	[]	[]	ĺ.		1	- 1+-21	[1	1	1	^	
	Learmont	h San V	Vito Sa	g Hill	Penticton	Penti	icton	Palehua	Pent	icton	Best set		
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245	28	24		-1	-1	-	-1 -1		-1		28		
410	52	54	R() 8	-1	-1		-1 -1		-1		52		
610	76	-1		-1	-1		1	-1		1	76		
1415	135	13	9	-1	-1	-	1	-1	1	1	135		
2695	192	19	0	-1	-1		1	-1	2	1	192		
2800	-1	-1	3	-1	-1	12	-1 -1		<u>,</u>	1	-1		
4995	211	27	9	-1	-1		1	-1		1	211		
8800	306	32	D	-1	-1	14	1	-1	<u></u>	1	306		
15400	568	58	588 -1		-1		1	-1		1	568	~	
un flux data	download	ed from : ft	p://ftp.swp	.noaa.gov,	/pub/lists/r	adio/rad	l.txt (-1	stands for	data noi	t available)			
	Cas	siopeia A	Cygnus A	Tauru	s A Sagit	arius A	Virgo A	30	273	Moon	Sur	i	
Flux	(Jy)	1947	1748	912	2. 5	520	223		42	718	1287	365	
oise Y-Facto	r (dB)	1.00	0.91	0.50) 0	.29	0.13		0.02	0.40	22.	35	
		Moon dis	tance (Km)	40292	9 New	Moon		Full Moon	-	New Moon			
		Moon	age (days)	4.41		6) ((60	Close		
Moon est	imated su	rface temp	erature (K)	238 3			-	\sim	-				

The predicted Noise Y-factor for the other objects is derived from the interpolation of the data published in the papers reported in the visual manual.

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Future function



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Future function : Noise Horizon



Noise Horizon at my site @ 1300 Mhz

The NOISE HORIZON function, working together with PstRotator, generates a "contour plot" that can be archived and then executed regularly to monitor changes at the observation site.

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Few more functions / improvements are in plan :

- \checkmark NOISE HORIZON with his own 3D analysis tool.
- ✓ Linearization of frequency domain plots.
- \checkmark Visibility improvements to operate under the sun.
- ✓ Introduce a new SDR : Air spy ... Adalm Pluto...
 B200.... need to define.

... any input and request will be much appreciated !

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Grazie !

Thank you !

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